

Clusters and the Cosmic X-ray Background

- (i) For a particular set of clusters the X-ray luminosity is found to be proportional to $T^{2.6}$, where T is the X-ray temperature. What is the temperature of a cluster whose luminosity is four times that of another cluster, whose temperature is $kT = 3$ keV?
- (ii) A particular cluster of galaxies has a gravitating or characteristic radius, $R_G = 2$ Mpc and a radial velocity dispersion (i.e. one-dimensional dispersion) of $\sigma_r = 1500$ km s $^{-1}$.
 - (a) Estimate the virial mass.
 - (b) Show that the escape velocity of a galaxy at R_G is $v_{\text{esc}} = \sqrt{6}\sigma_r$.
- (iii) Explain why type 1 quasars *cannot* account for most of the hard cosmic X-ray background. What were the principal technological advances necessary to discover what kind of sources *do* constitute the hard X-ray background?

Cosmology

- (ii) Explain the principal characteristics of the three types of Universe resulting from the solution of Einstein's field equations for a simple expanding Universe, with zero cosmological constant.
- (iii) Calculate the critical density of the Universe, $\rho_c = 3H_0^2/8\pi G$, in units of g cm $^{-3}$, using $H_0 = 50$ km s $^{-1}$ Mpc $^{-1}$.
- (iv) The energy density of the Cosmic Microwave Background, u_{CMB} , at the current epoch is 0.26 eV cm $^{-3}$. Convert this to an equivalent mass density in units of g cm $^{-3}$ and calculate the ratio of the CMB radiation density to the critical density. Can the CMB close the Universe?
- (v) What is the role of inflation in the Big Bang theory?